

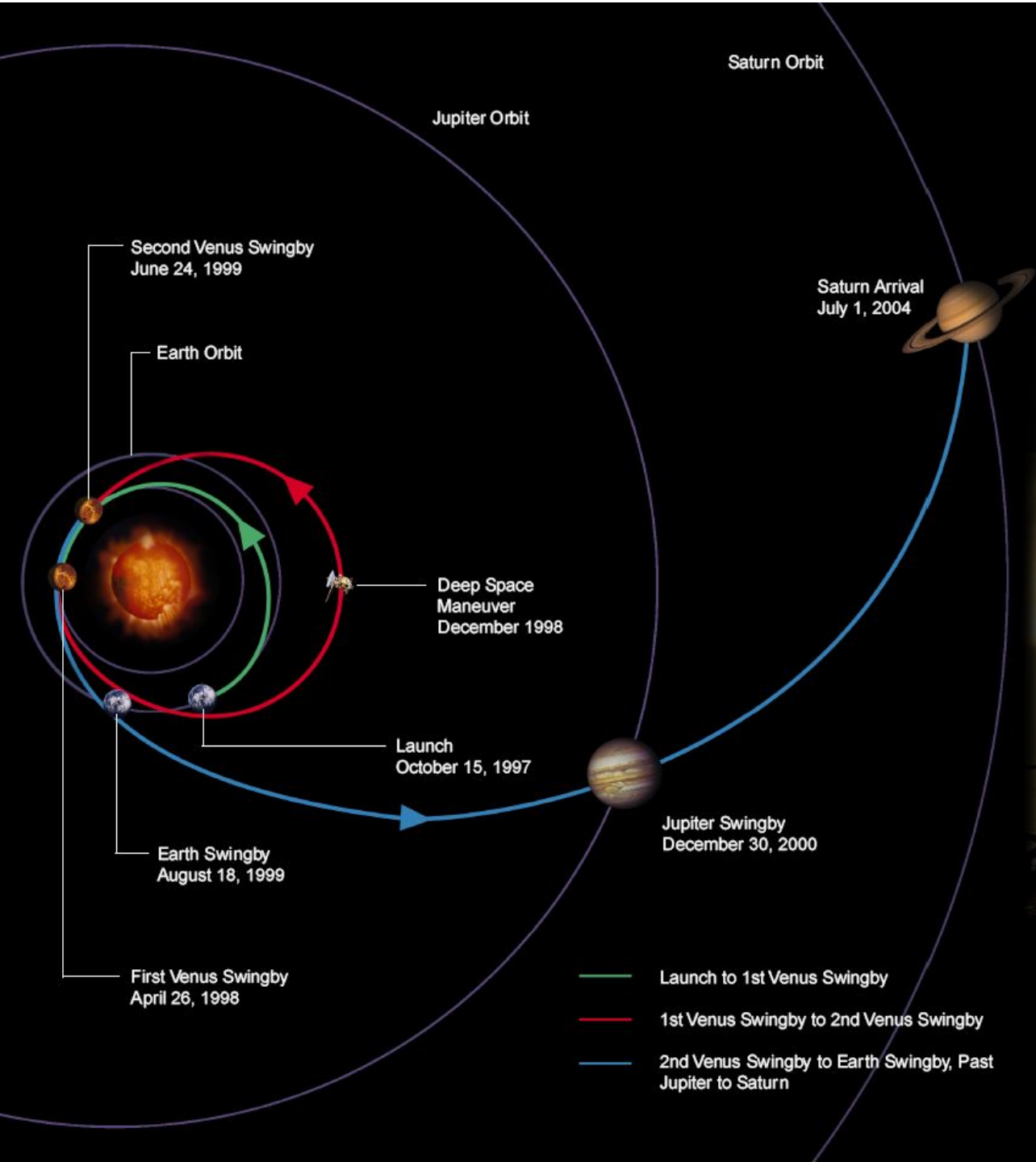
PROFONDITÀ DI UN MARE EXTRATERRESTRE



Marco Mastrogiuseppe

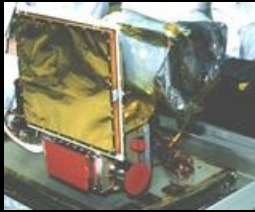
MISSIONE CASSINI – HUYGENS: GRAN FINALE

15 Febbraio 2018
Università La Sapienza
Roma



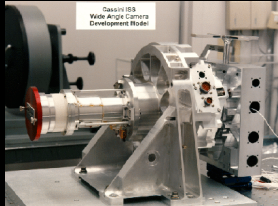
Radar & Radio Science

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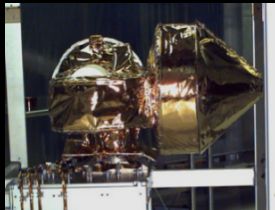
✗ Non è possibile visualizzare questa

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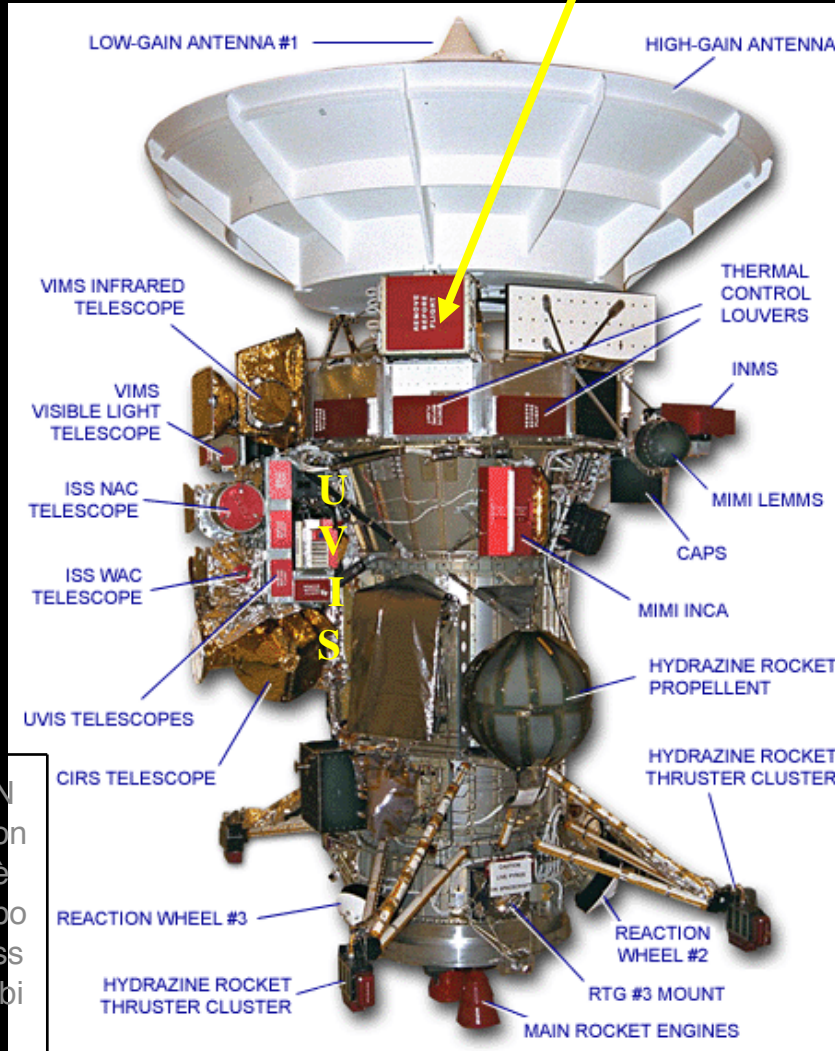
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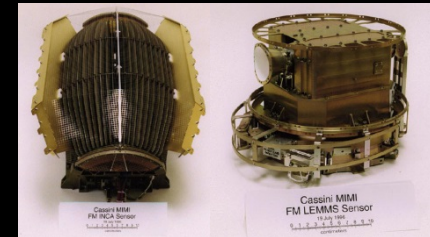
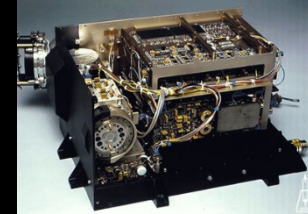
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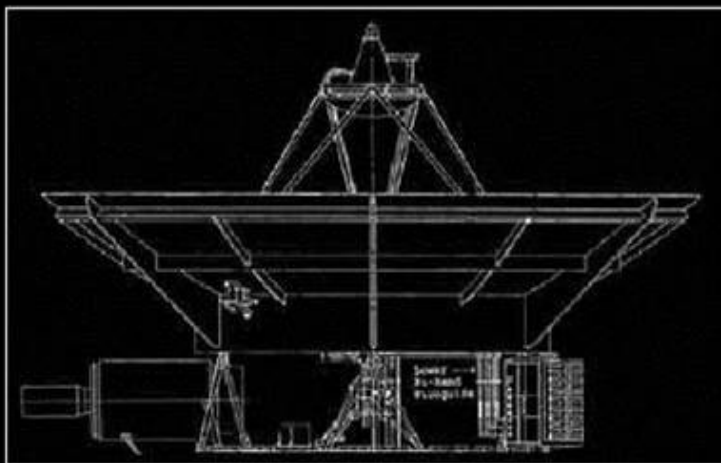


MIMI INCA & LEMMS

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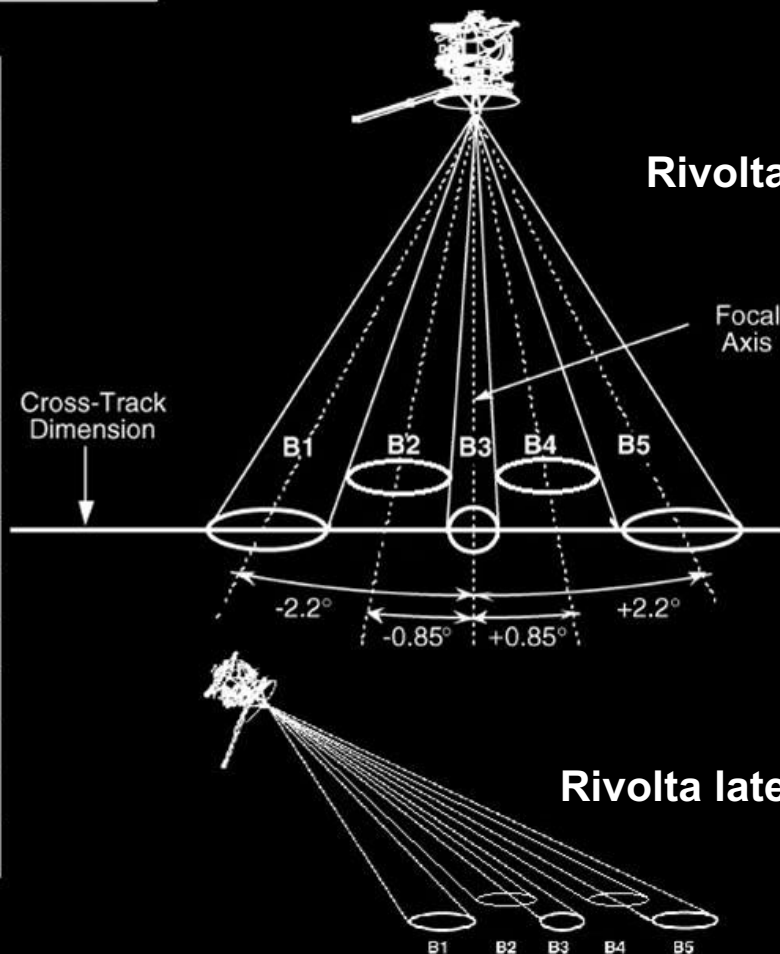
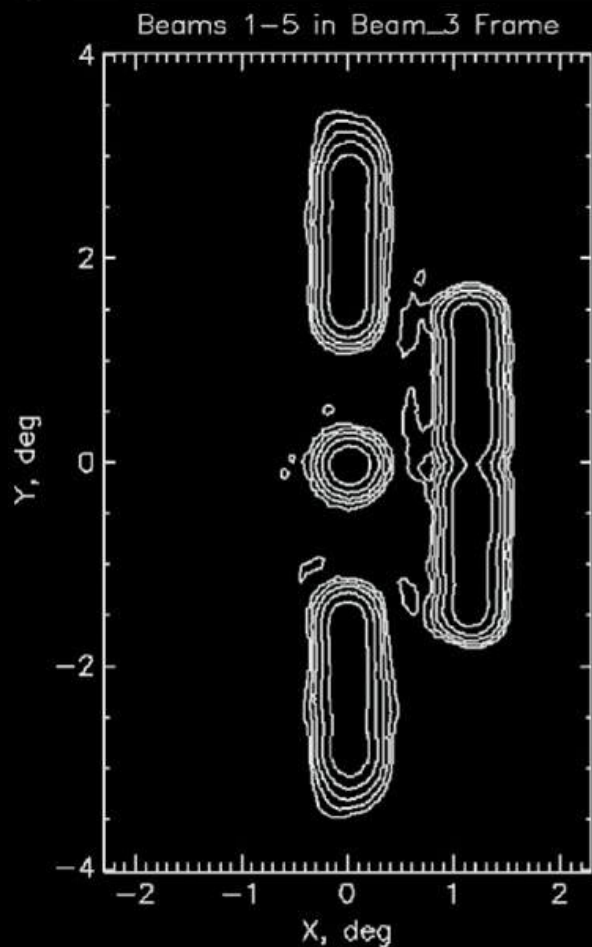


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High Gain Antenna

ASI ha selezionato Alenia Aerospazio a Roma come 'contractor' sia per l'high gain antenna (HGA) che per il RADAR Radio Frequency Electronics Subsystem (RFES).



Rivolta al nadir - Altimetrie

Rivolta lateralmente - Immagini

Cassini: A Multimode RADAR

- **Hi-Res SAR**

Altitude 1000-1500 km, Resolution 380 (azim.) x 600 (range) m

- **Low-Res SAR**

Altitude 1500-4000 km, Resolution 600 x 2500 m

- **Hi-Res Altimetry**

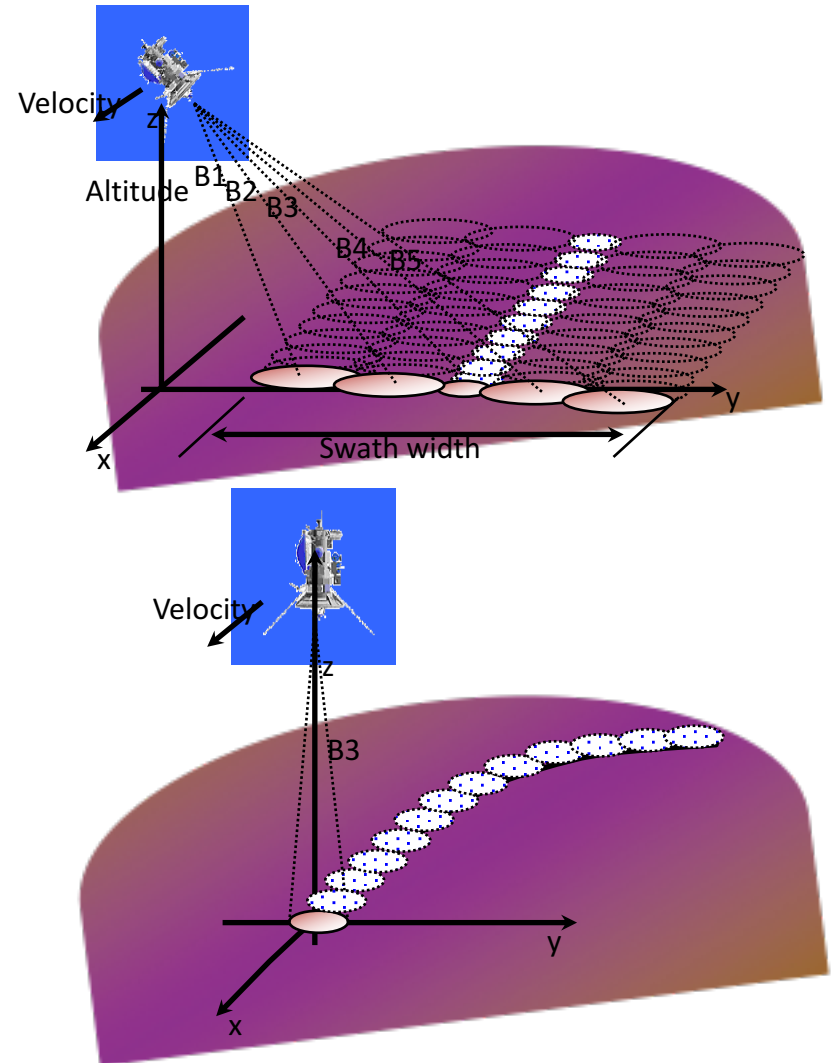
Altitude 4000-9000 km, Footprint few tens of km, Vertical resolution 35 m

- **Low-Res Altimetry (Scatterometry)**

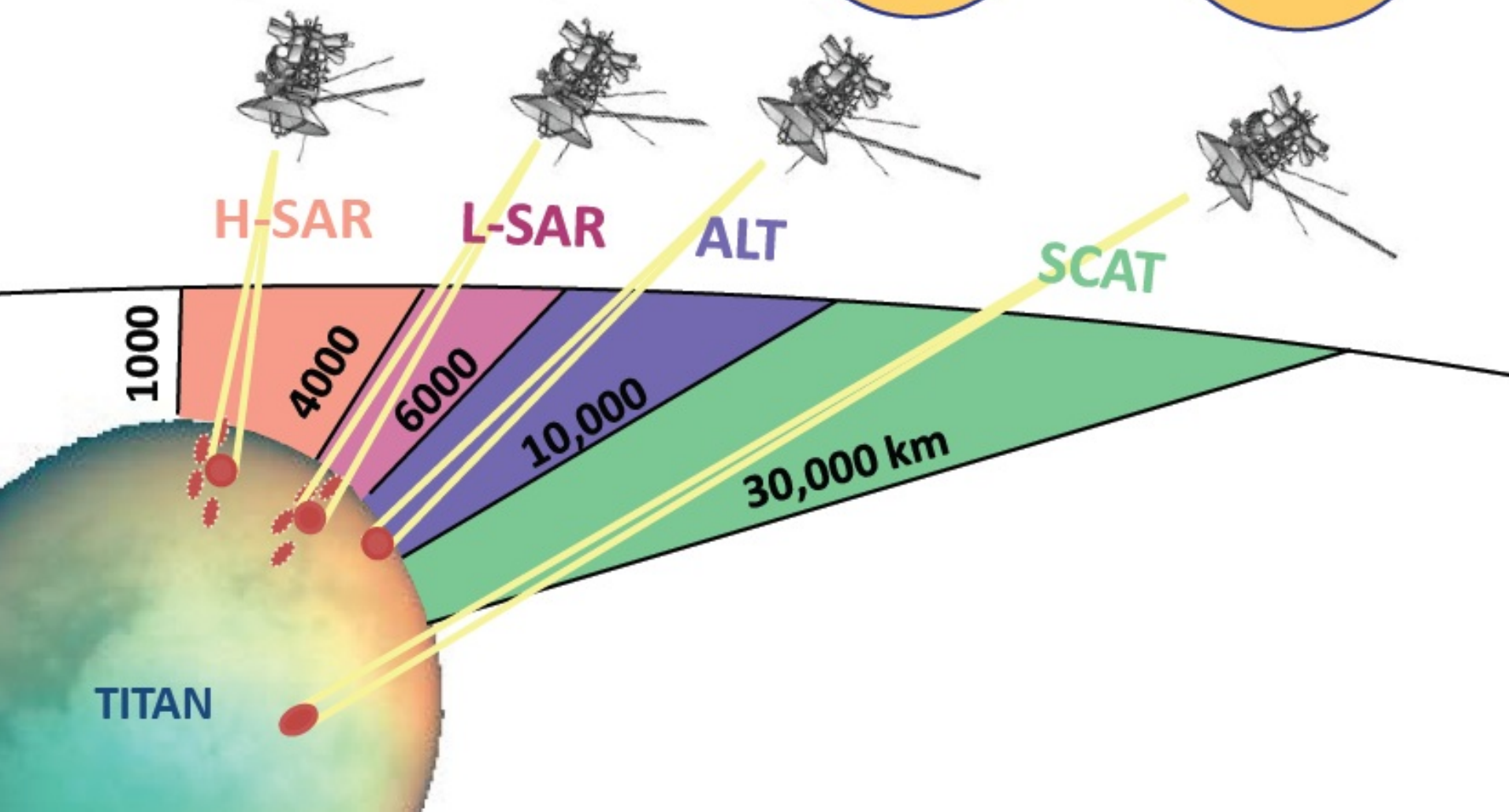
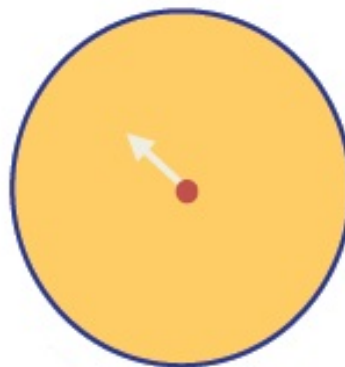
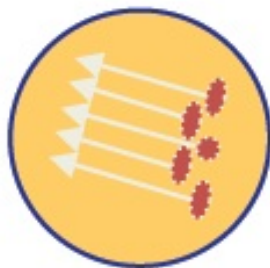
Altitude 9000-22500 km, Footprint hundreds of km, Vertical resolution 1415 m

- **Radiometer**

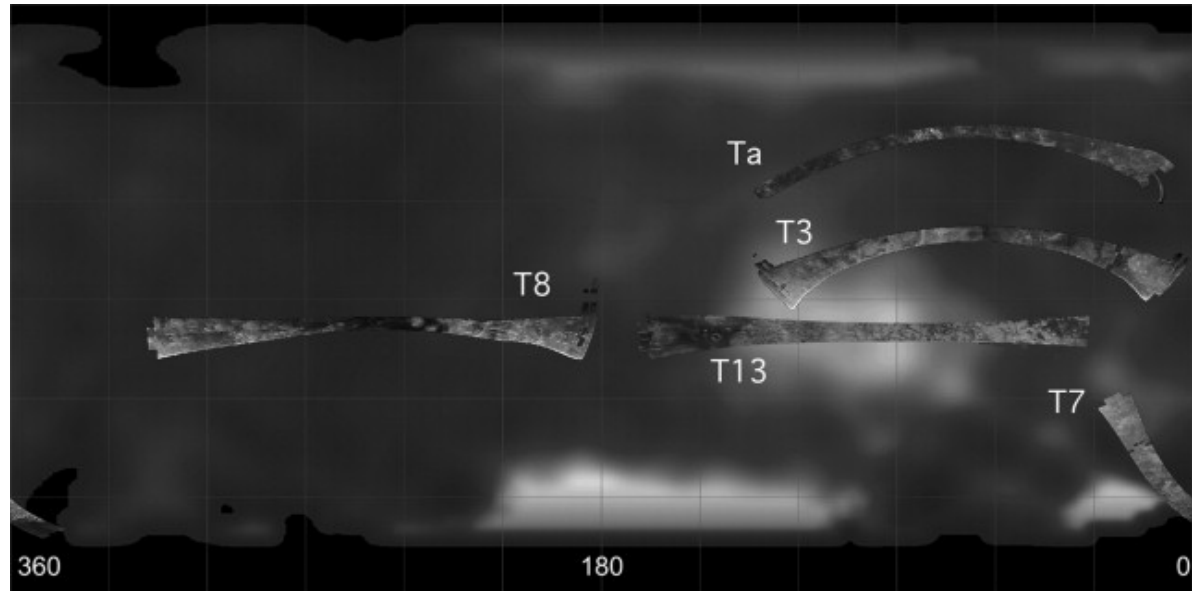
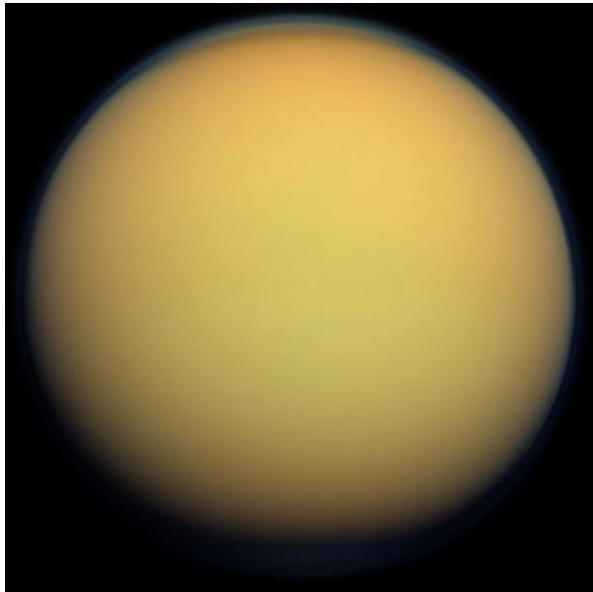
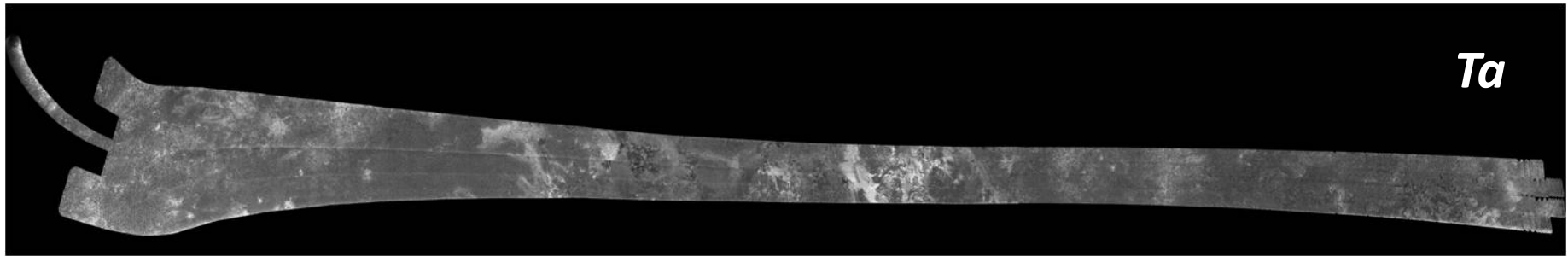
Altitude 1000-100000 km, 1-sec 1- σ noise = 0.025 K, Half power beamwidth 0.35 deg



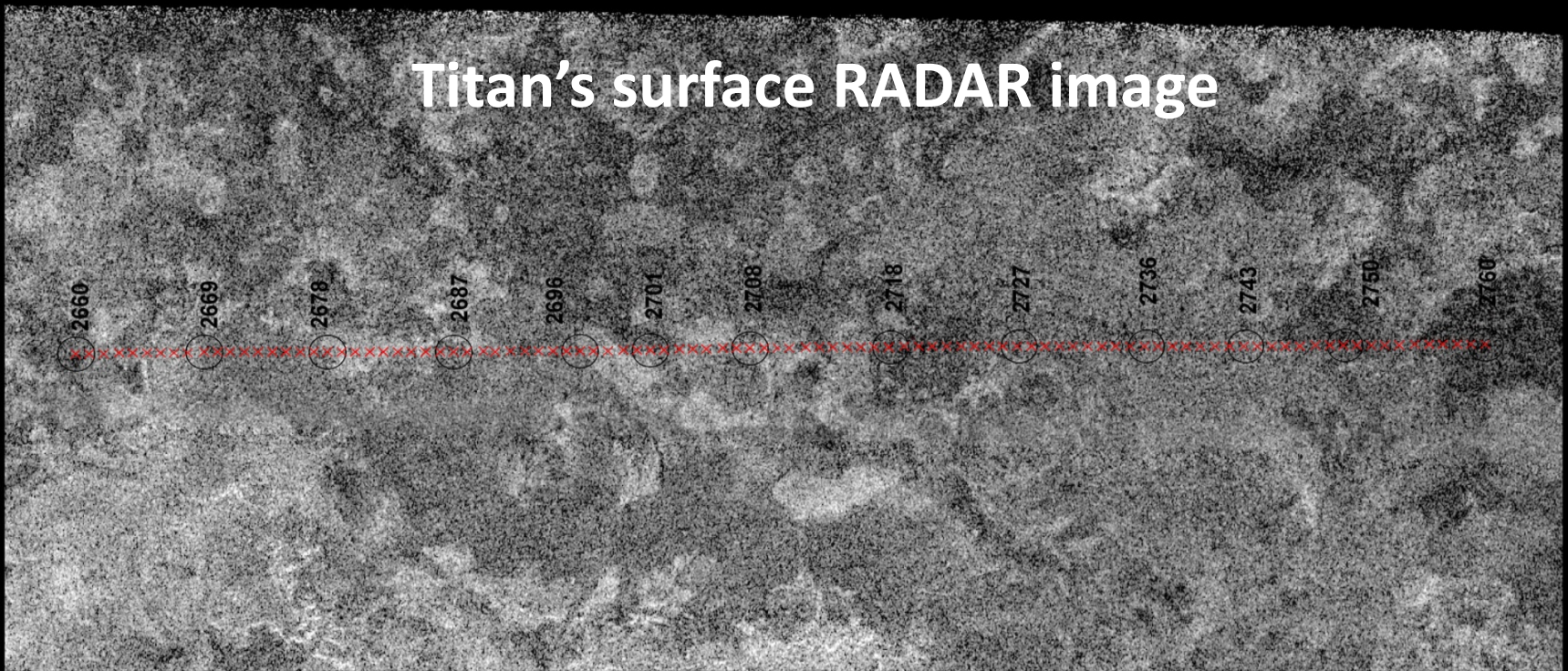
Surface Tracks



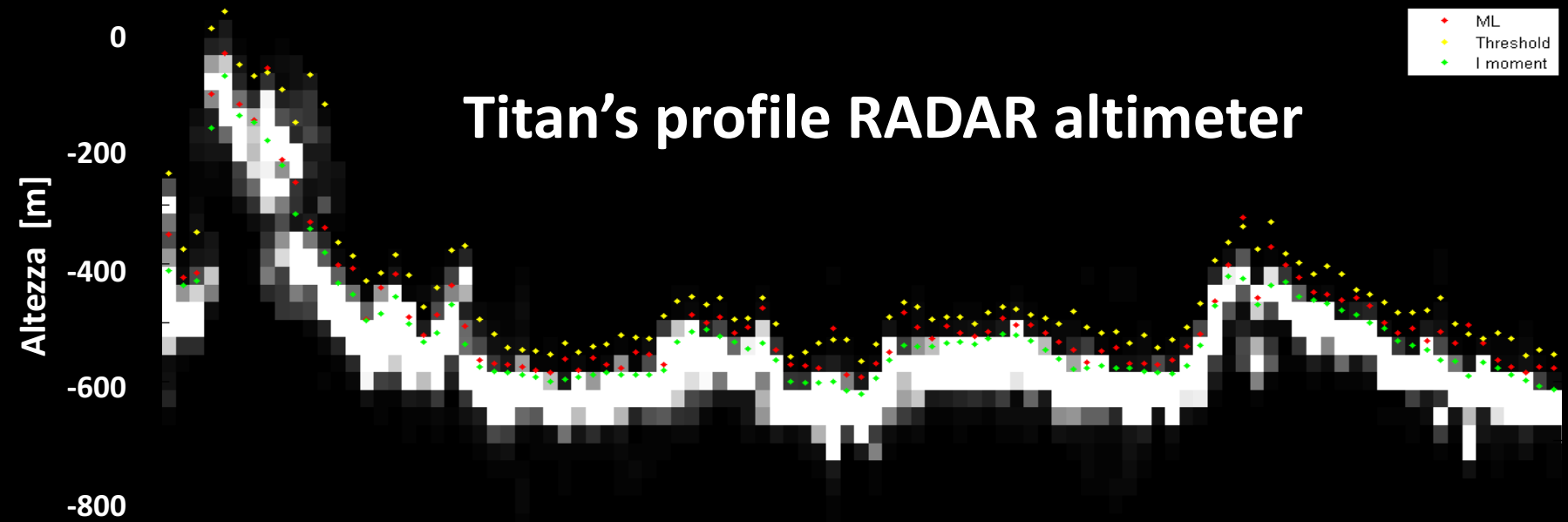
Ta: First RADAR image of Titan



Titan's surface RADAR image

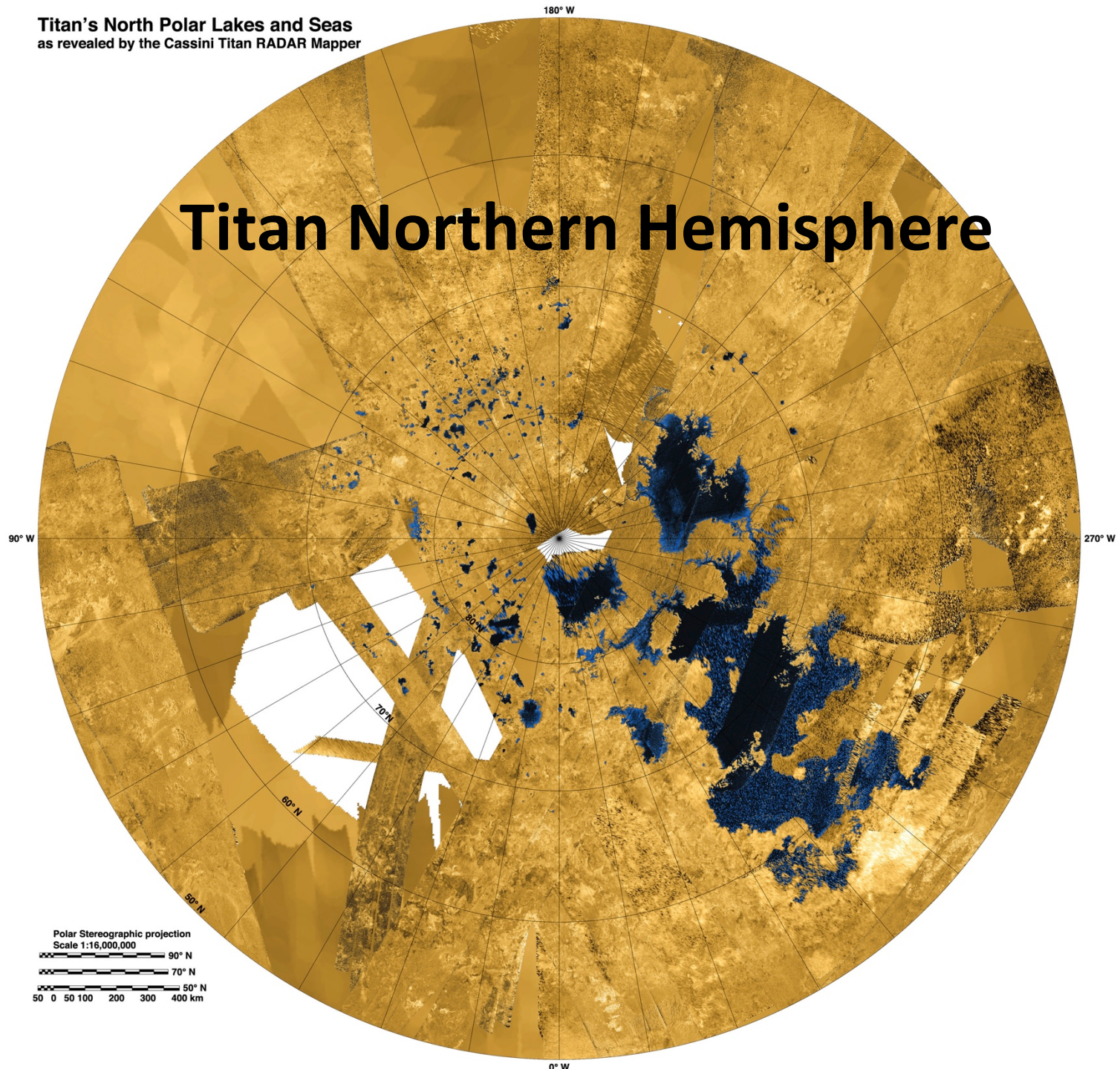


Titan's profile RADAR altimeter

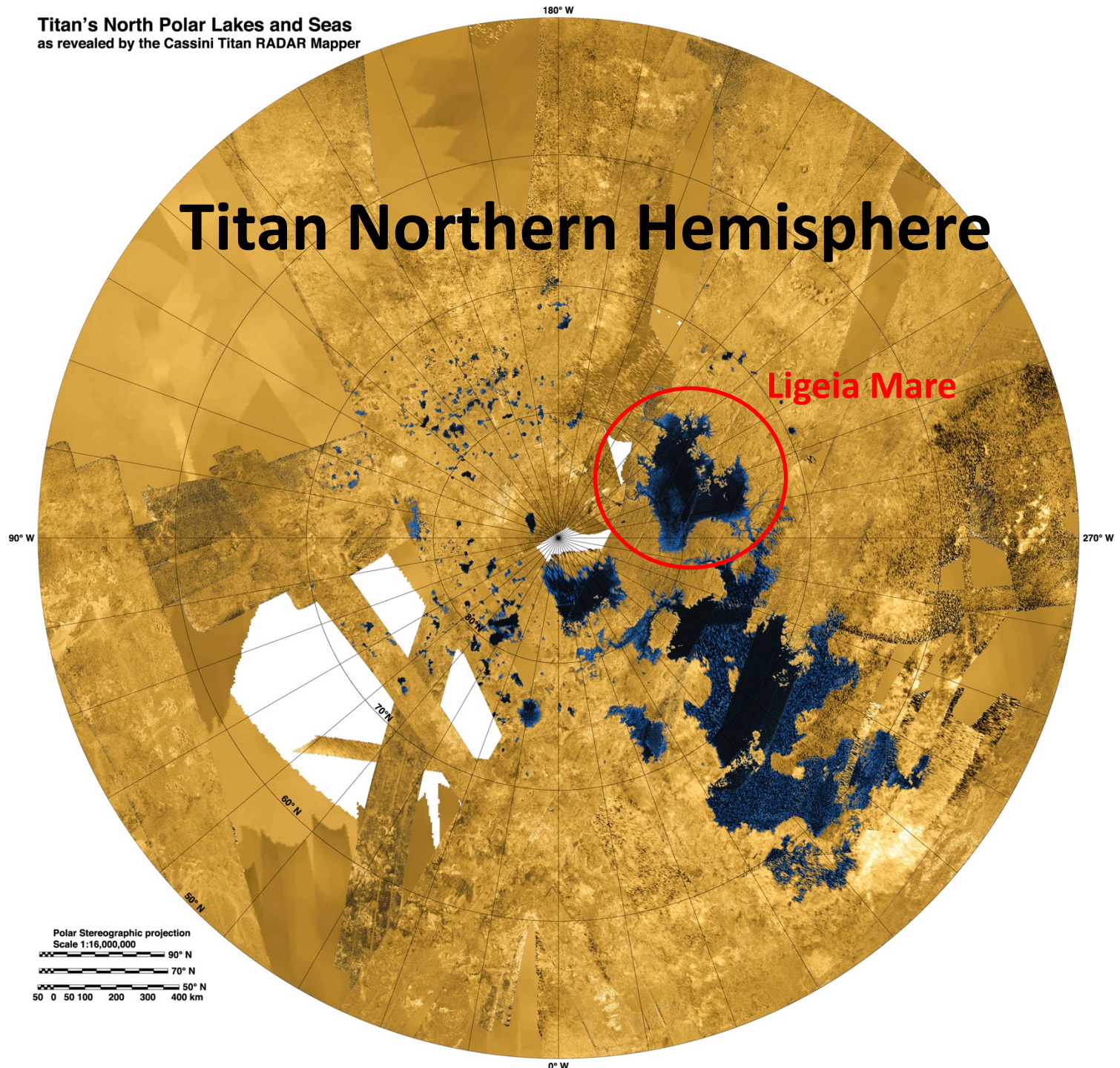


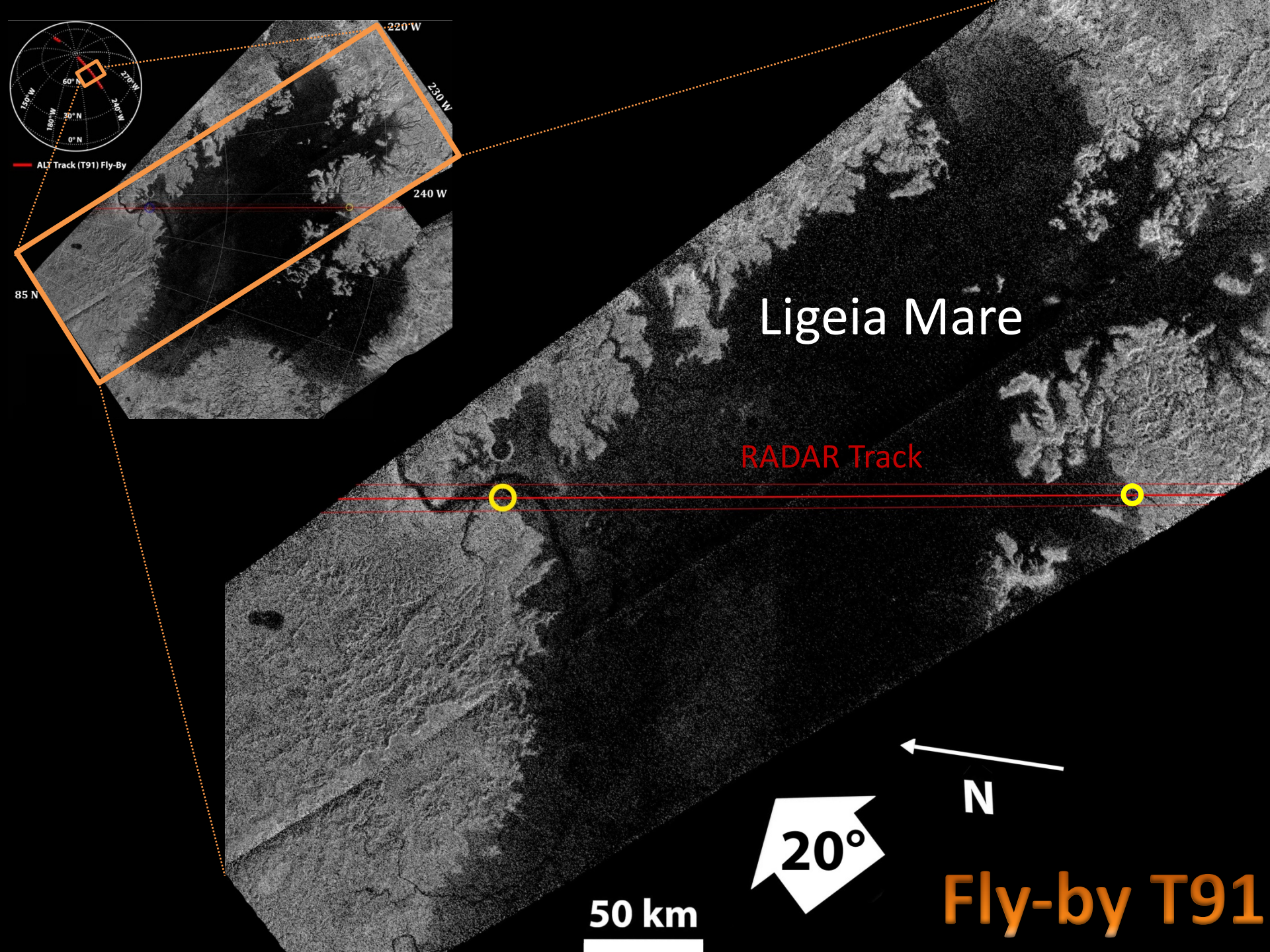
Titan's North Polar Lakes and Seas
as revealed by the Cassini Titan RADAR Mapper

Titan Northern Hemisphere

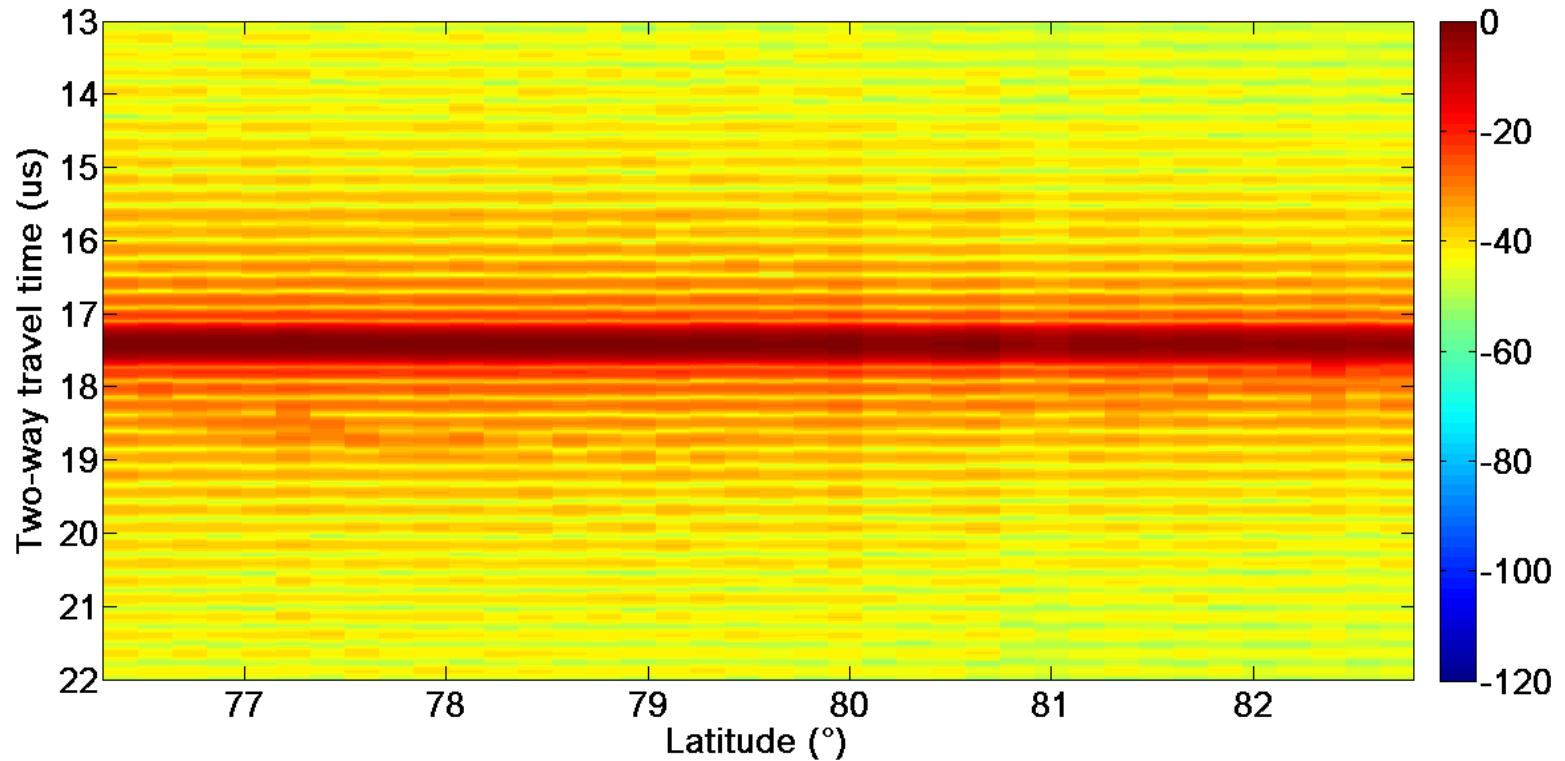


Titan Northern Hemisphere



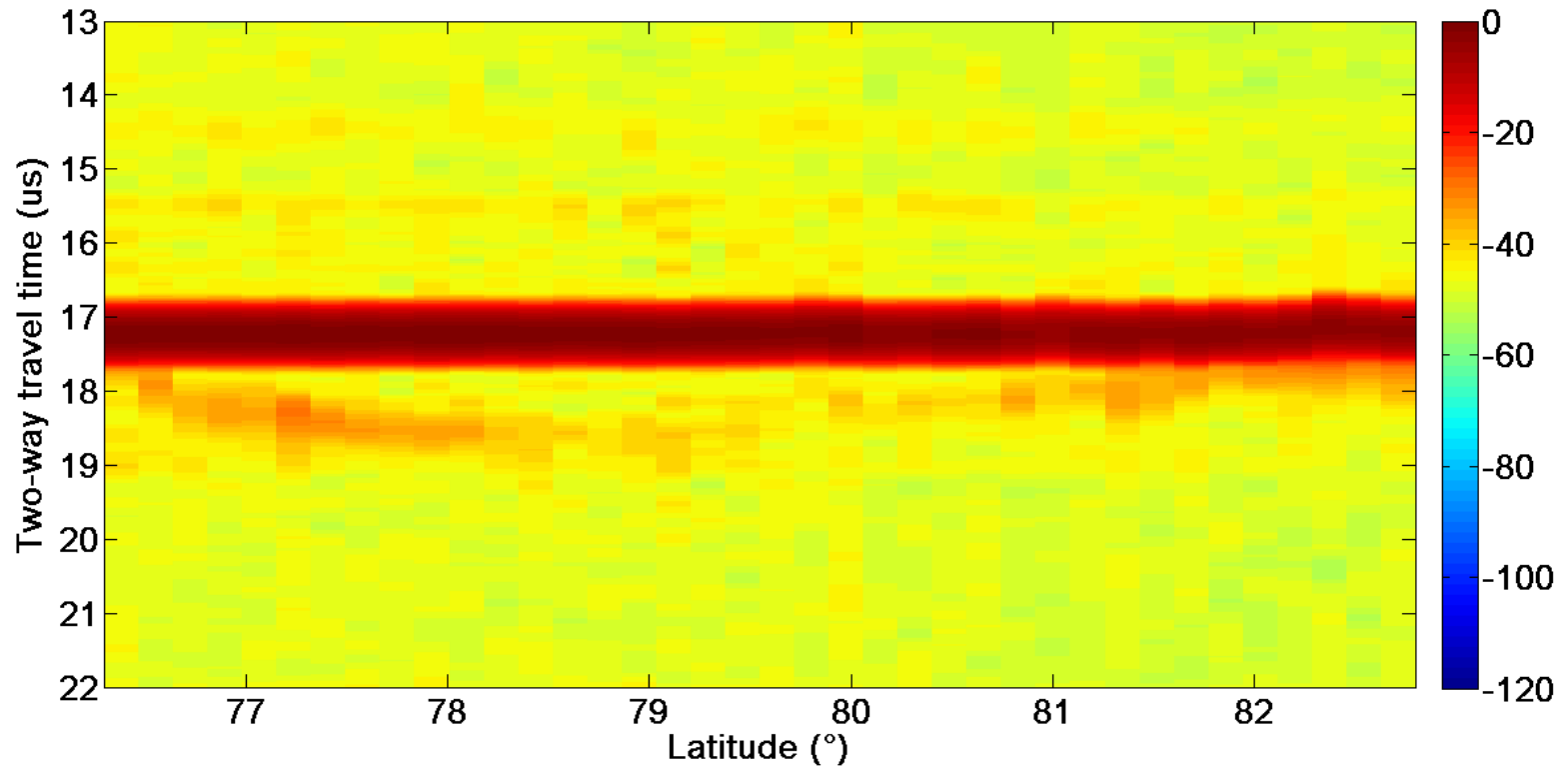


Data Processing



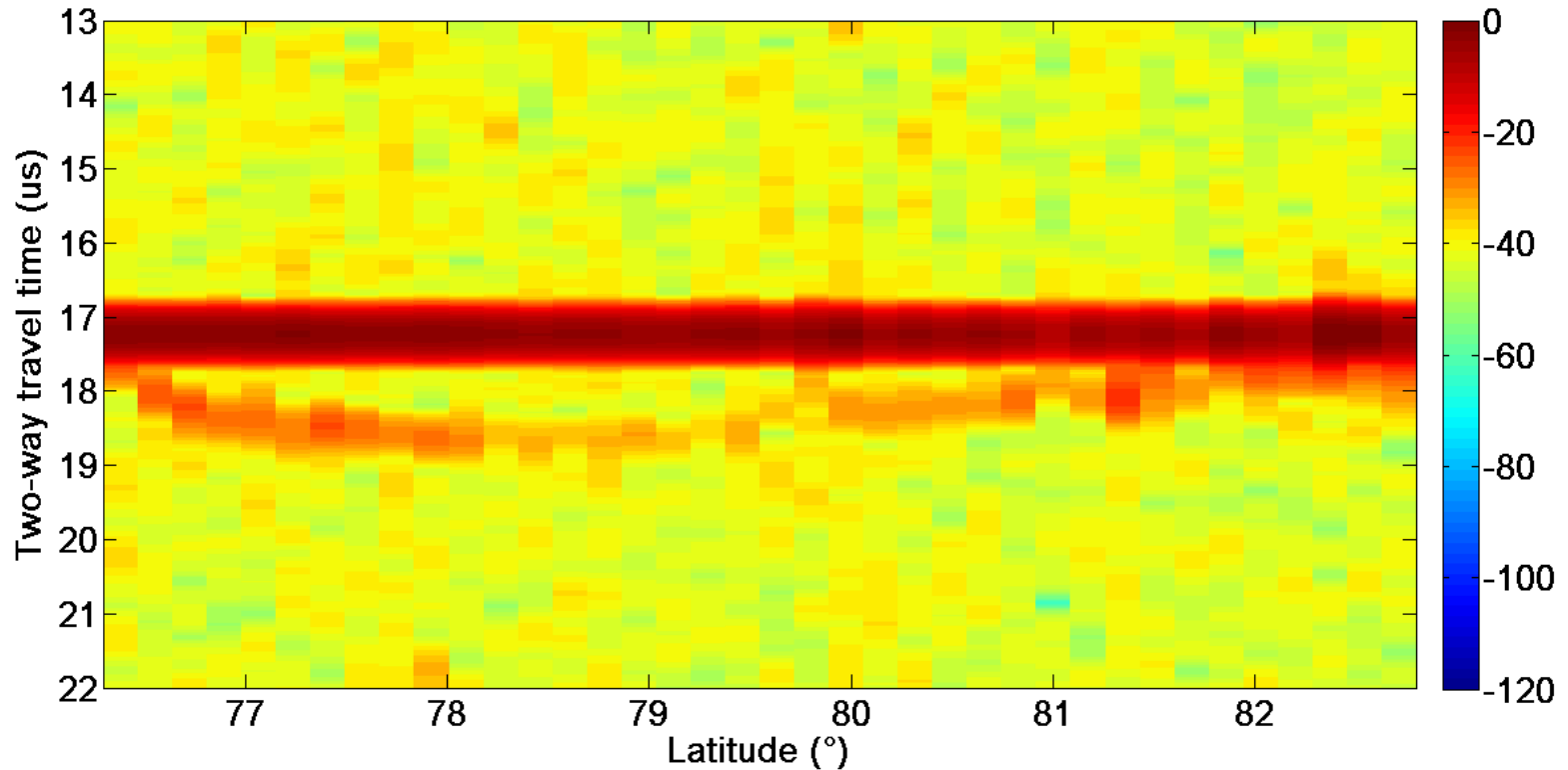
1 step processing

Data Processing



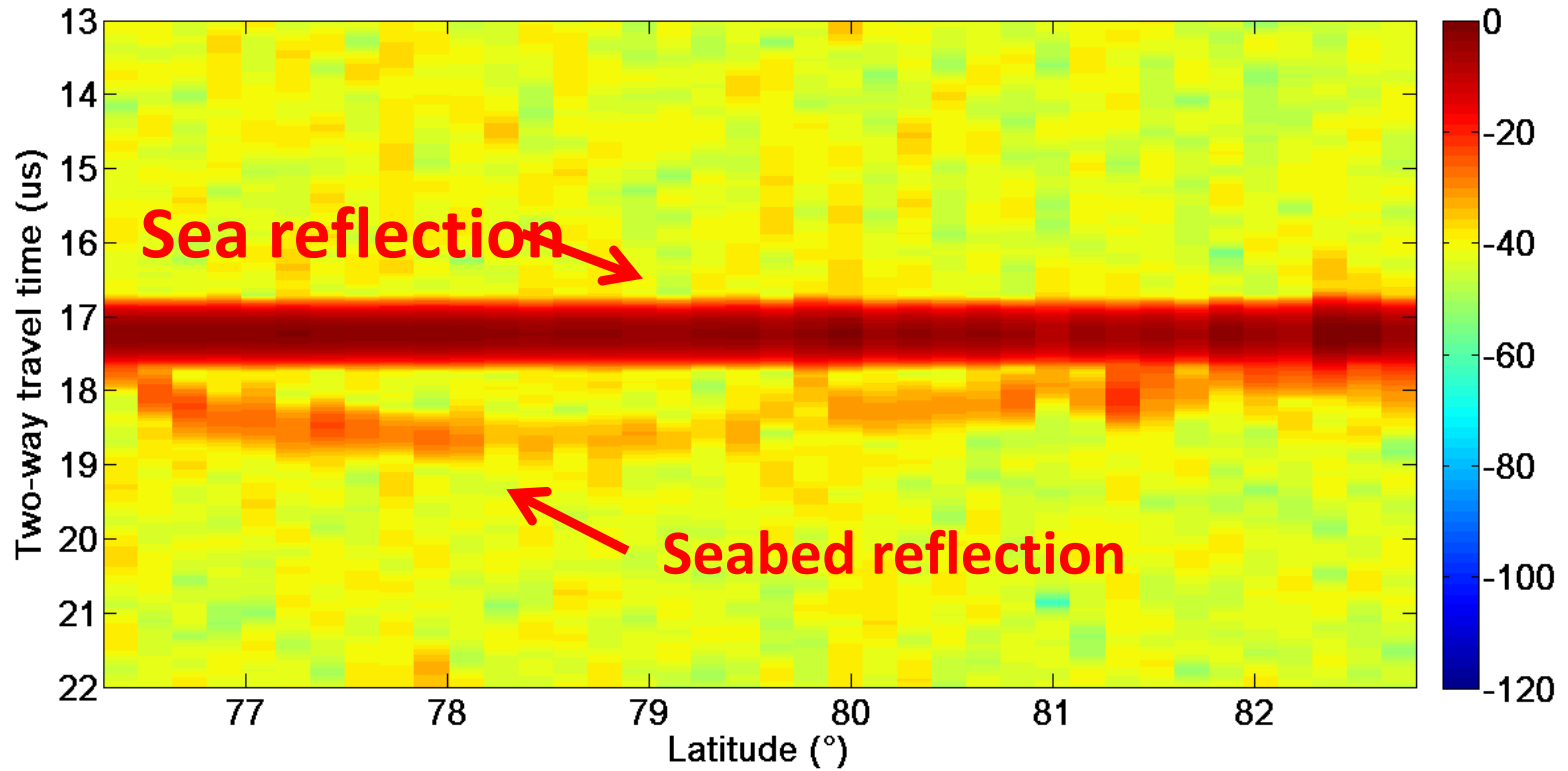
II step processing

Data Processing

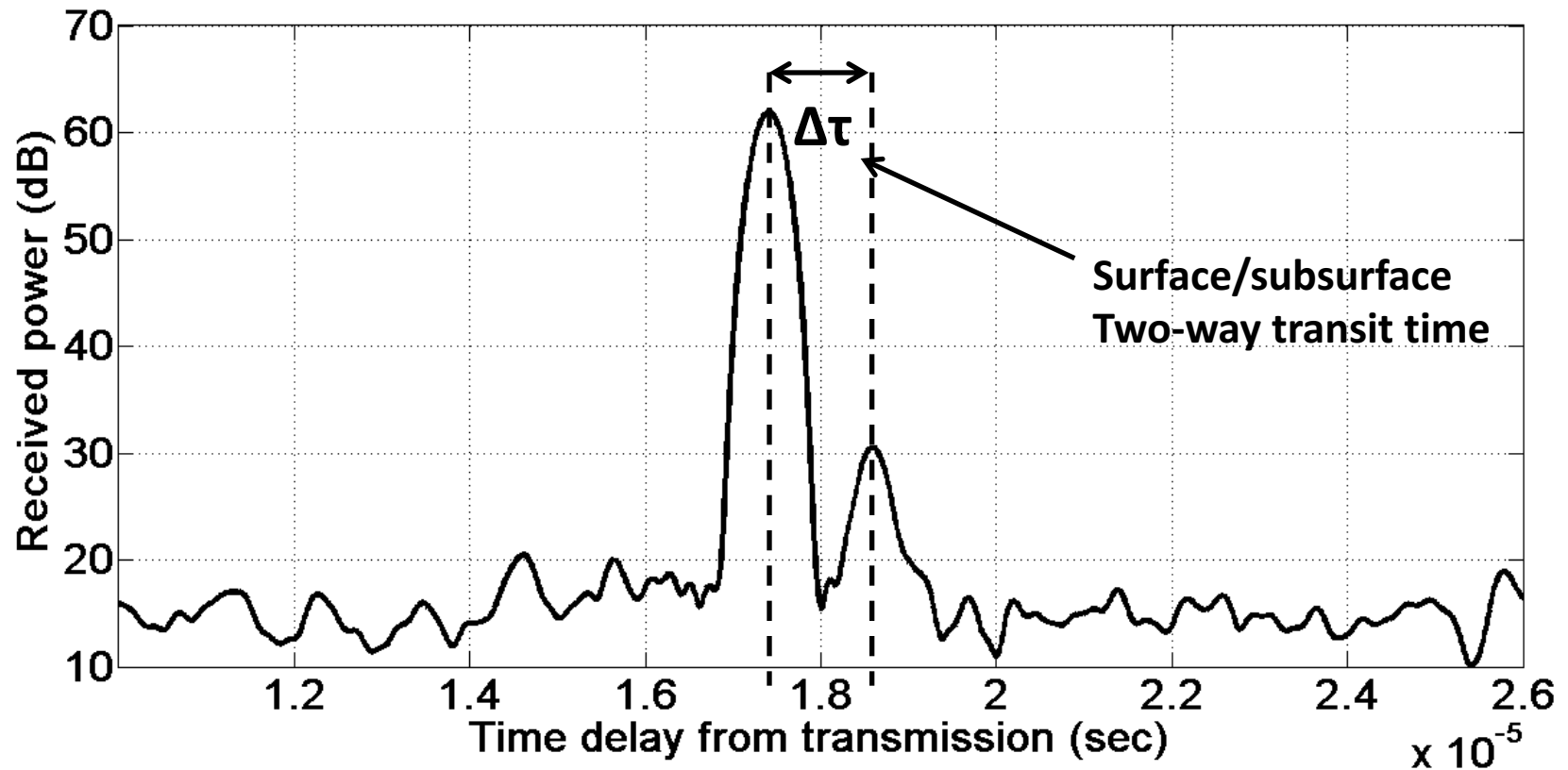


III step processing

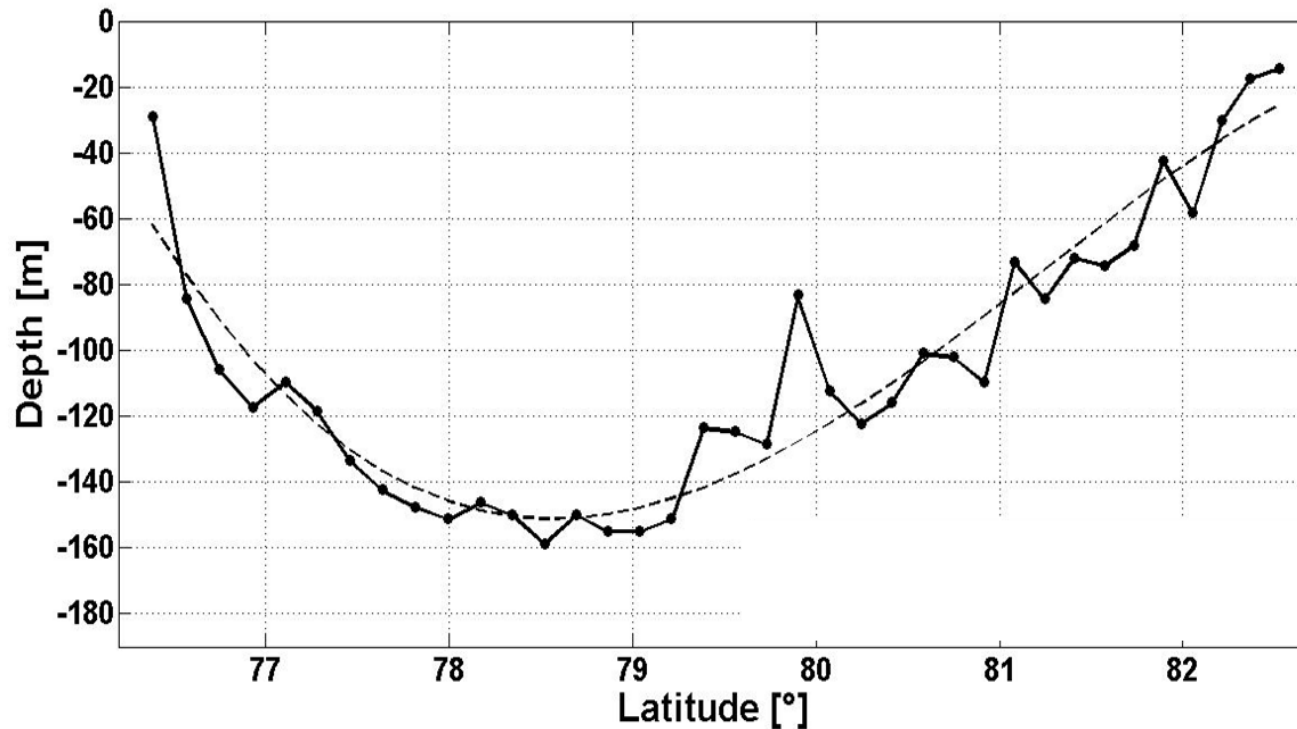
First radargram of Ligeia



Depth Measurements



Ligeia Mare Bathymetry

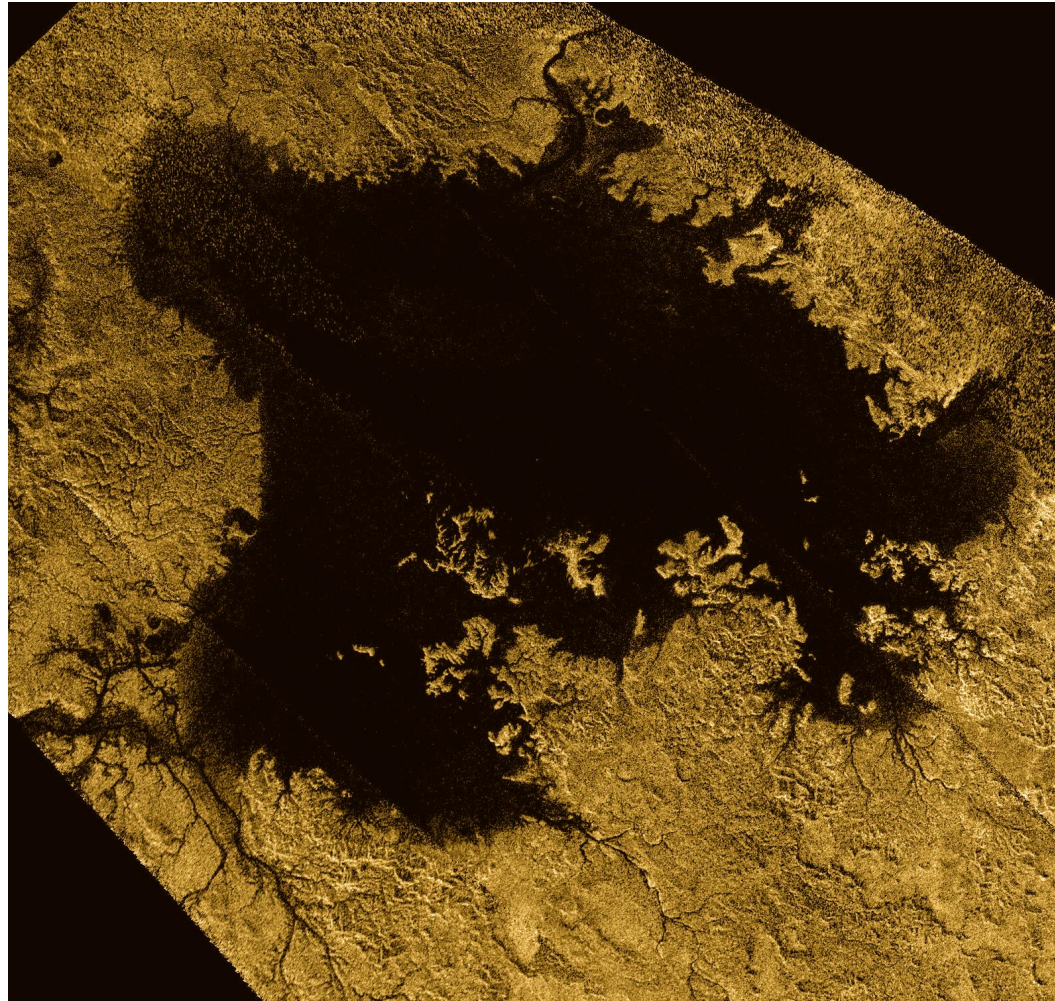


Maximum depth 160m

Implications: Total volume

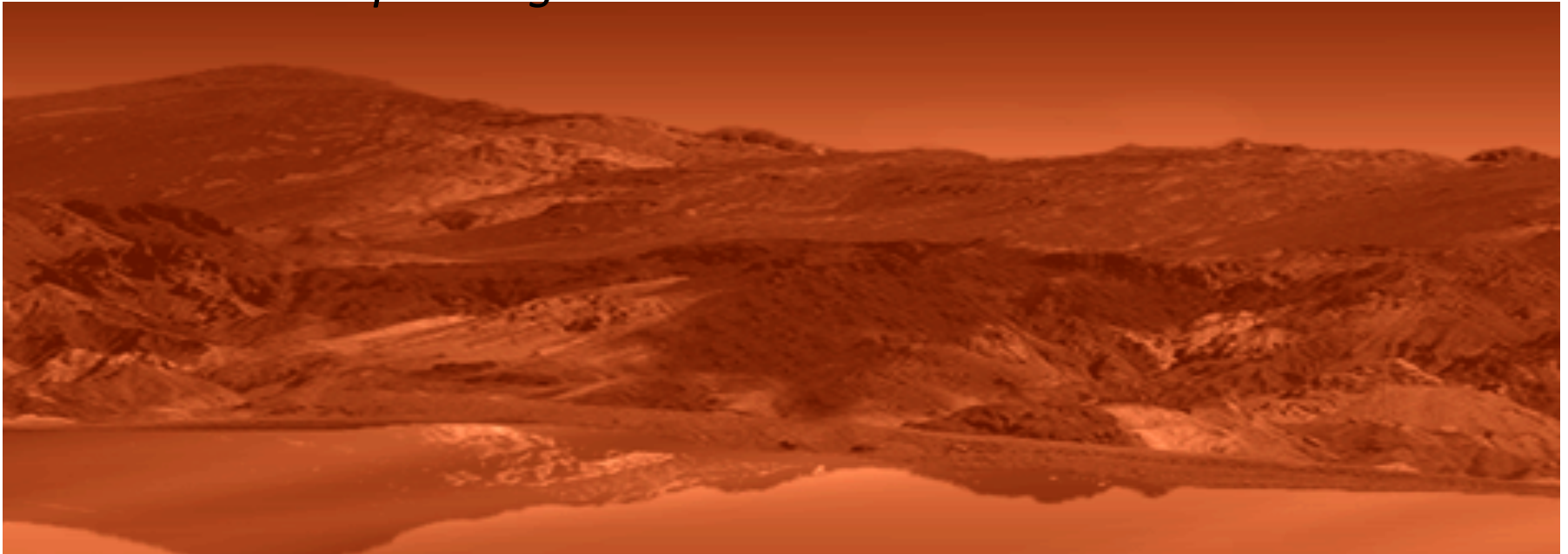
*Extrapolated total liquid volume
of Ligeia Mare: **about 3,500
cubic miles (14,000 cubic
kilometers) OR 8,000
gigatons of carbon***

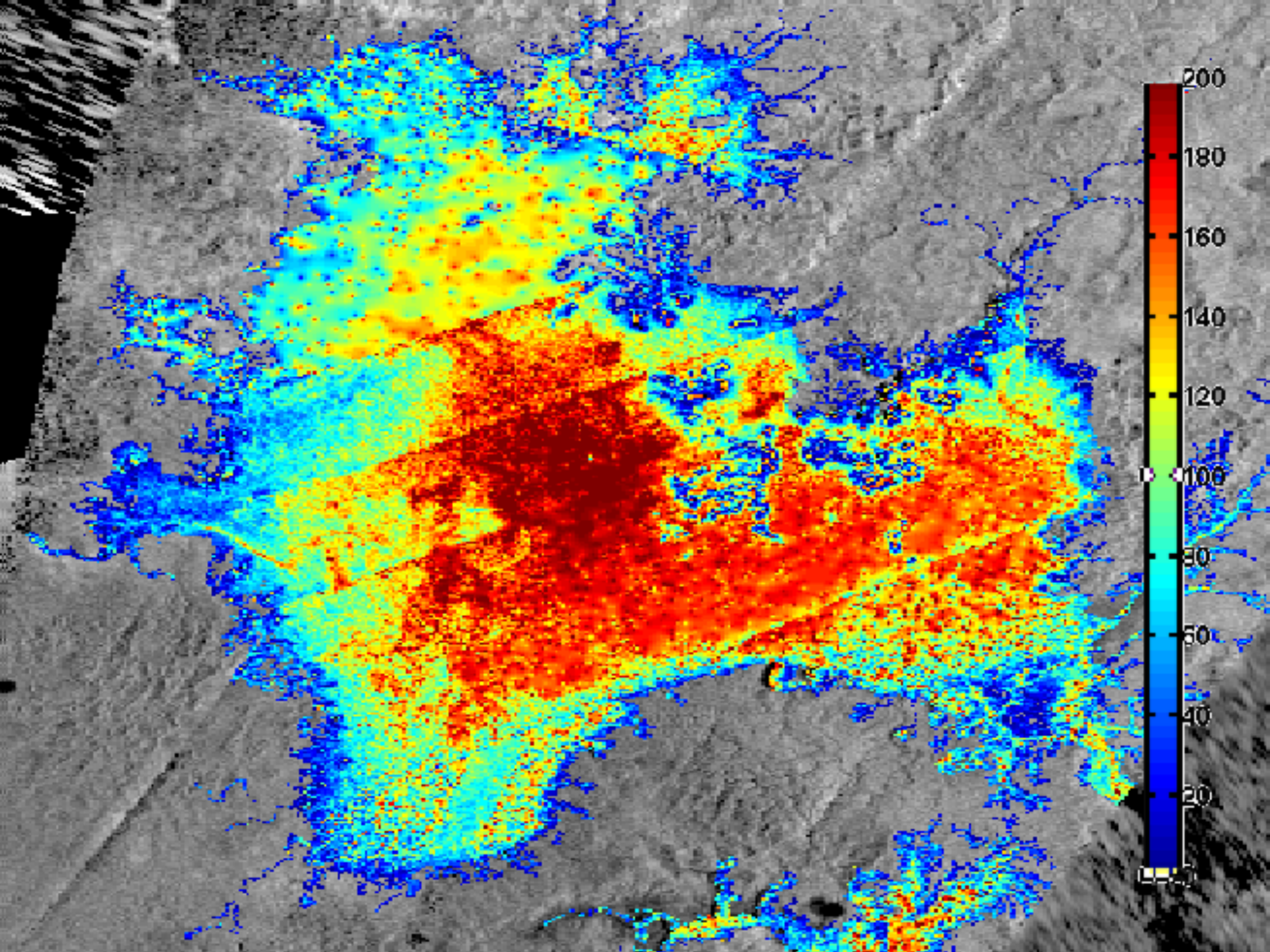
*About 55 times the proven oil
reserves on Earth !*



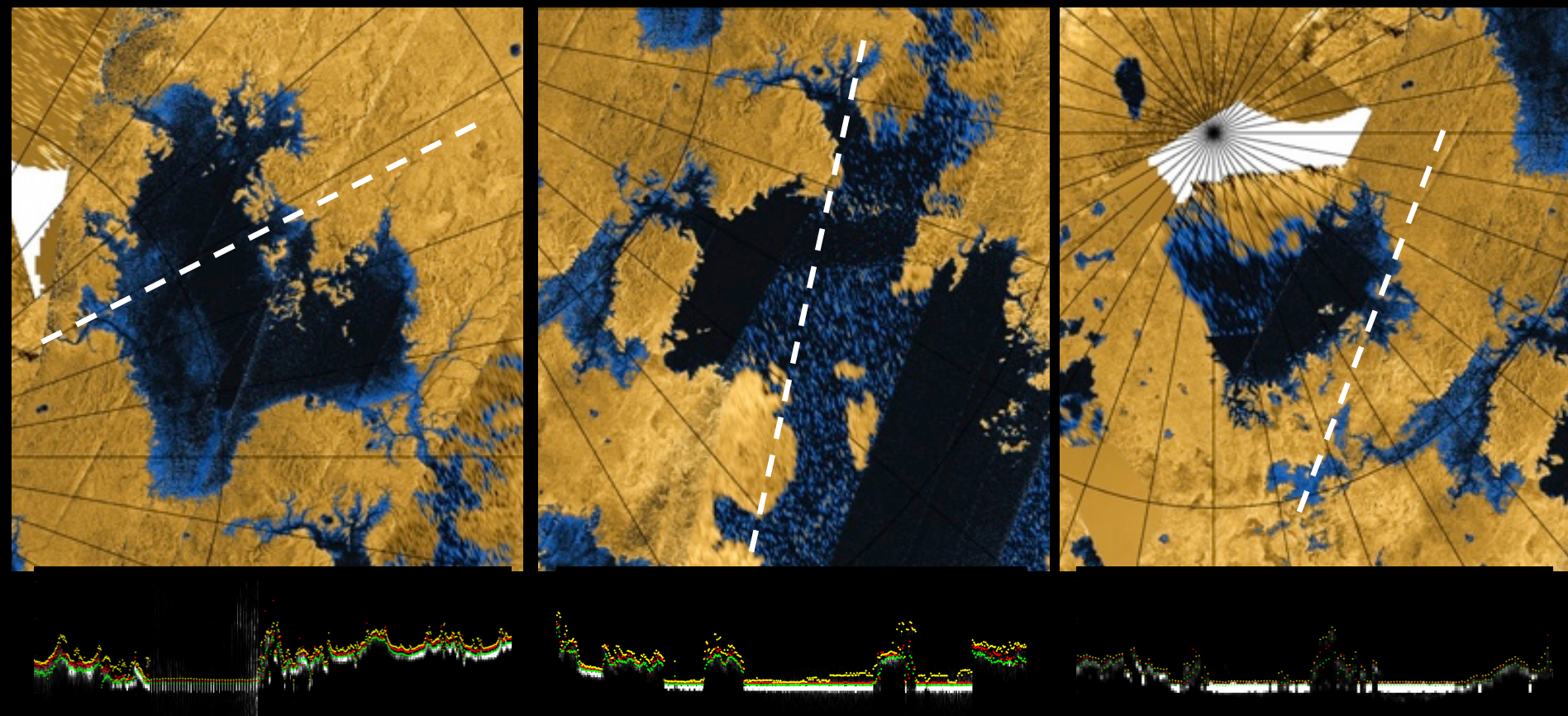
Liquid Composition

- *The analysis of Cassini radio signal attenuation through the sea demonstrates that the liquid is remarkably transparent, requiring under Titan conditions a composition that is **nearly pure methane-ethane**.*
- *Molecular nitrogen, which will be dissolved in the liquid at the 5-15% level depending on the methane-ethane ratio.*

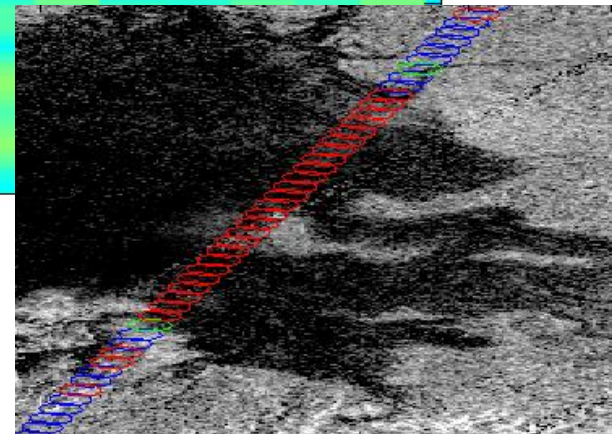
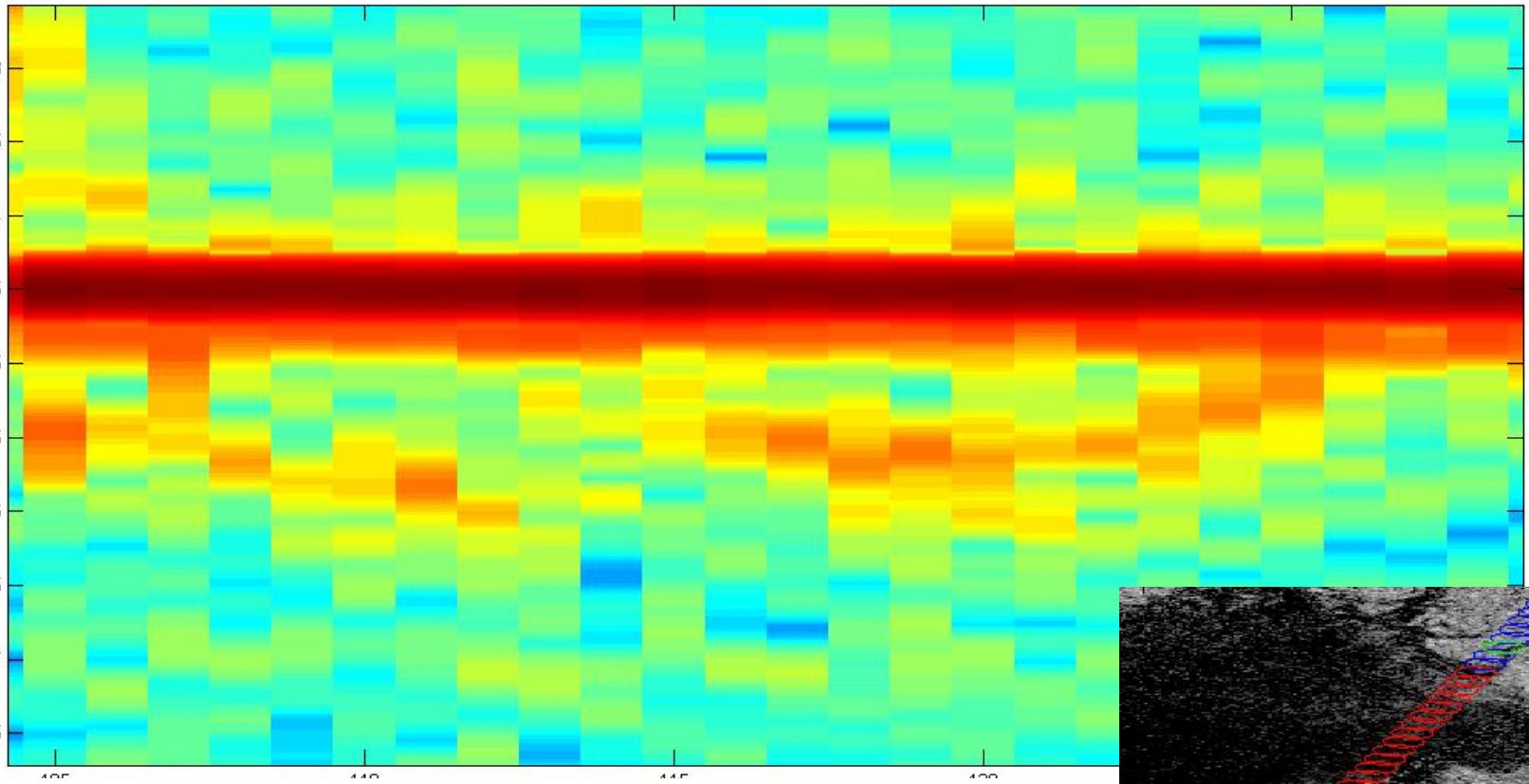




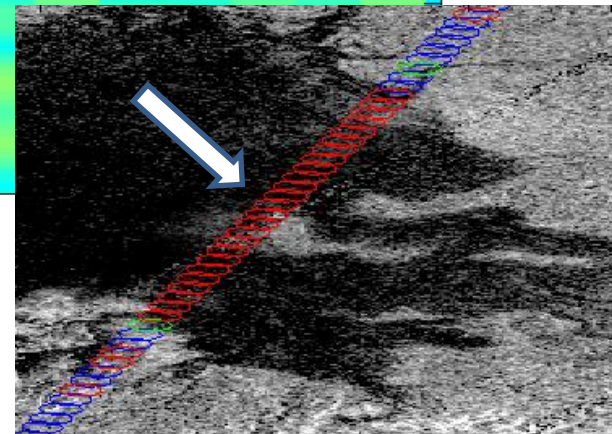
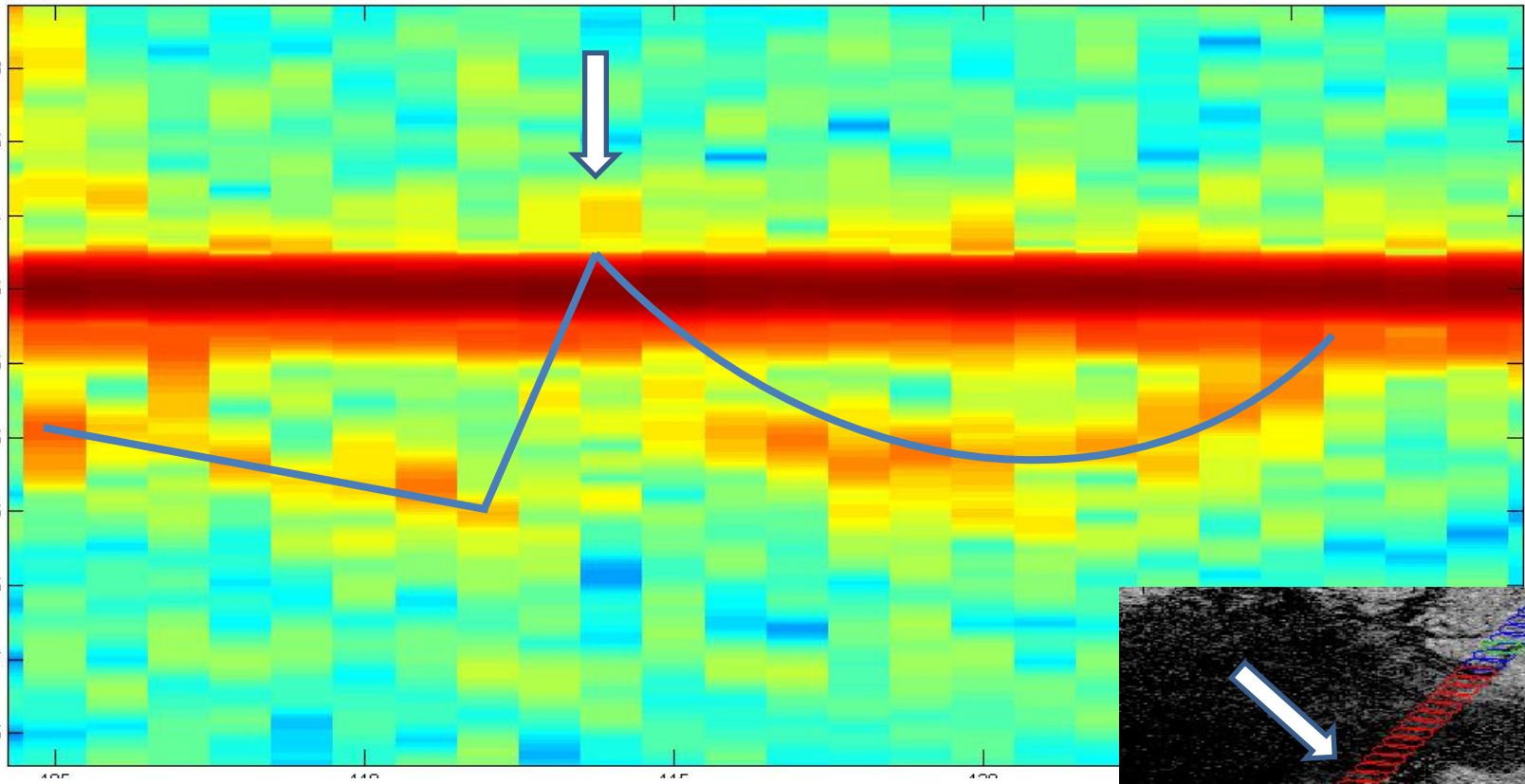
Bathymetry and Composition of Titan's Northern lakes from the Cassini RADAR Altimeter



T108 Punga Mare

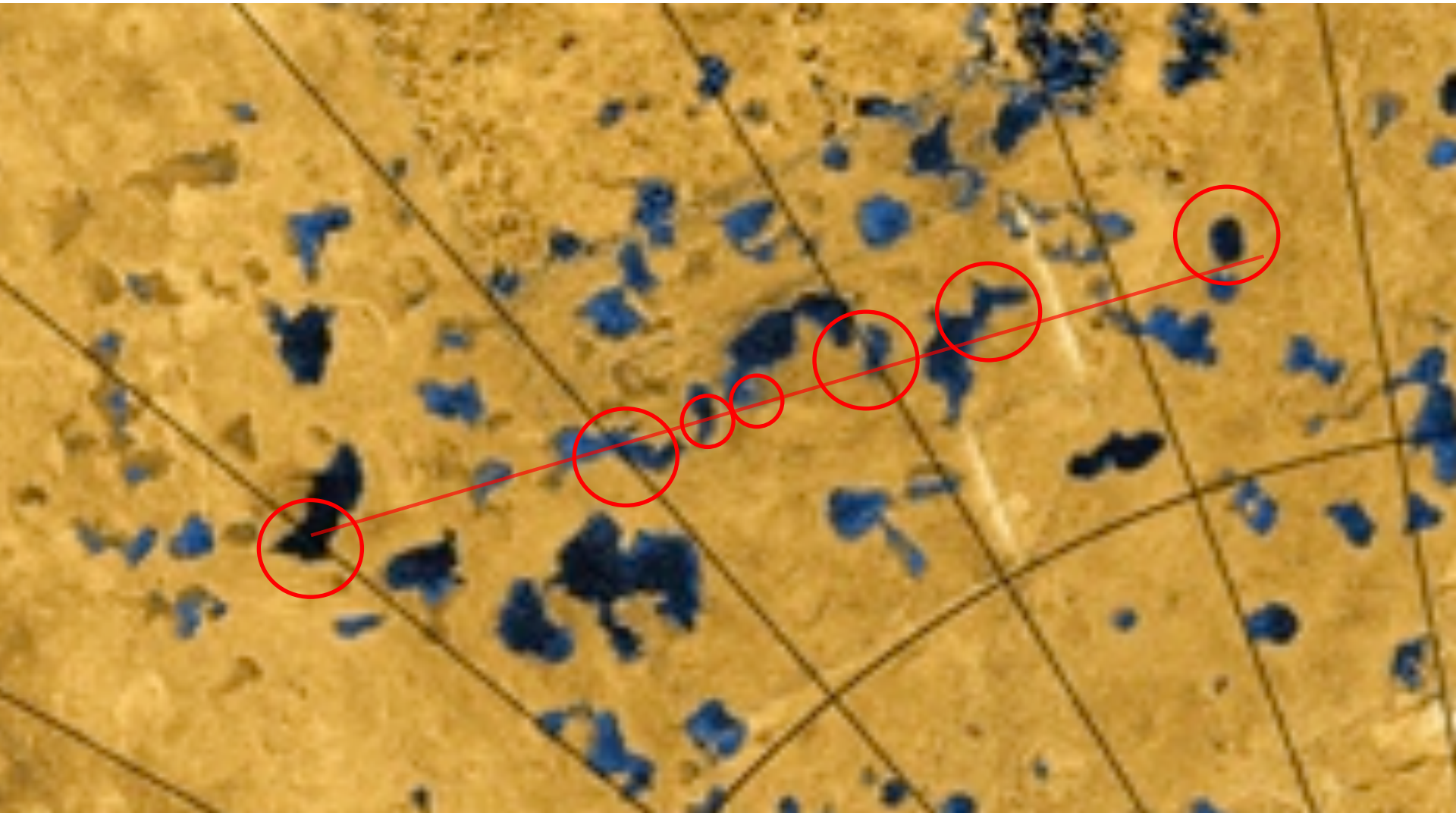


T108 Punga Mare

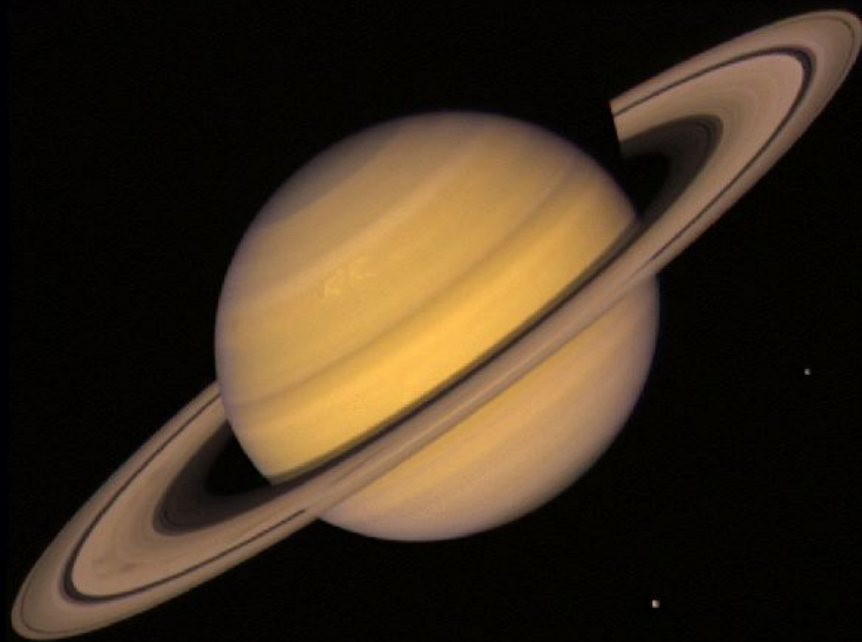


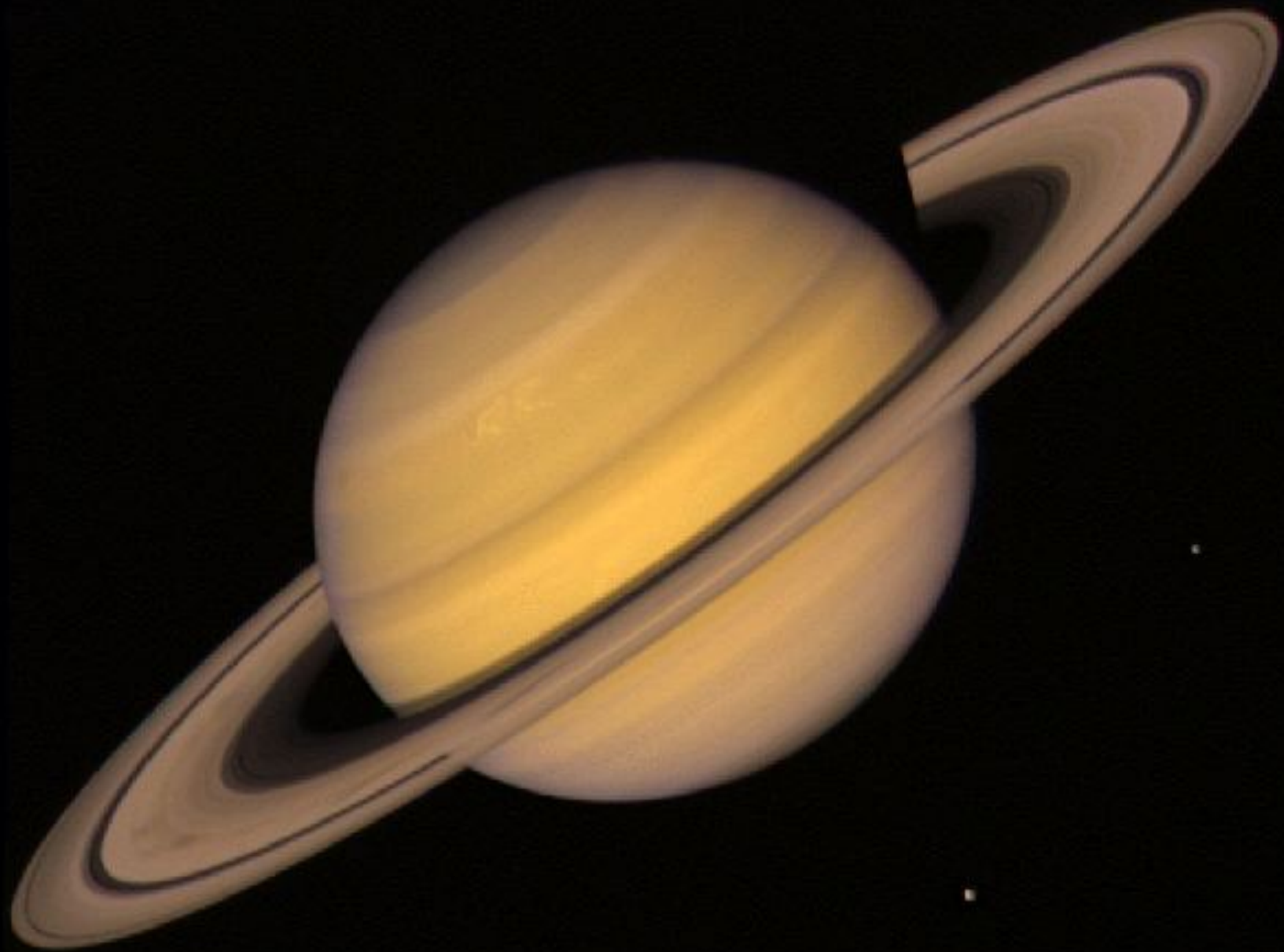
T126 Fly-by (April 22nd 2017)

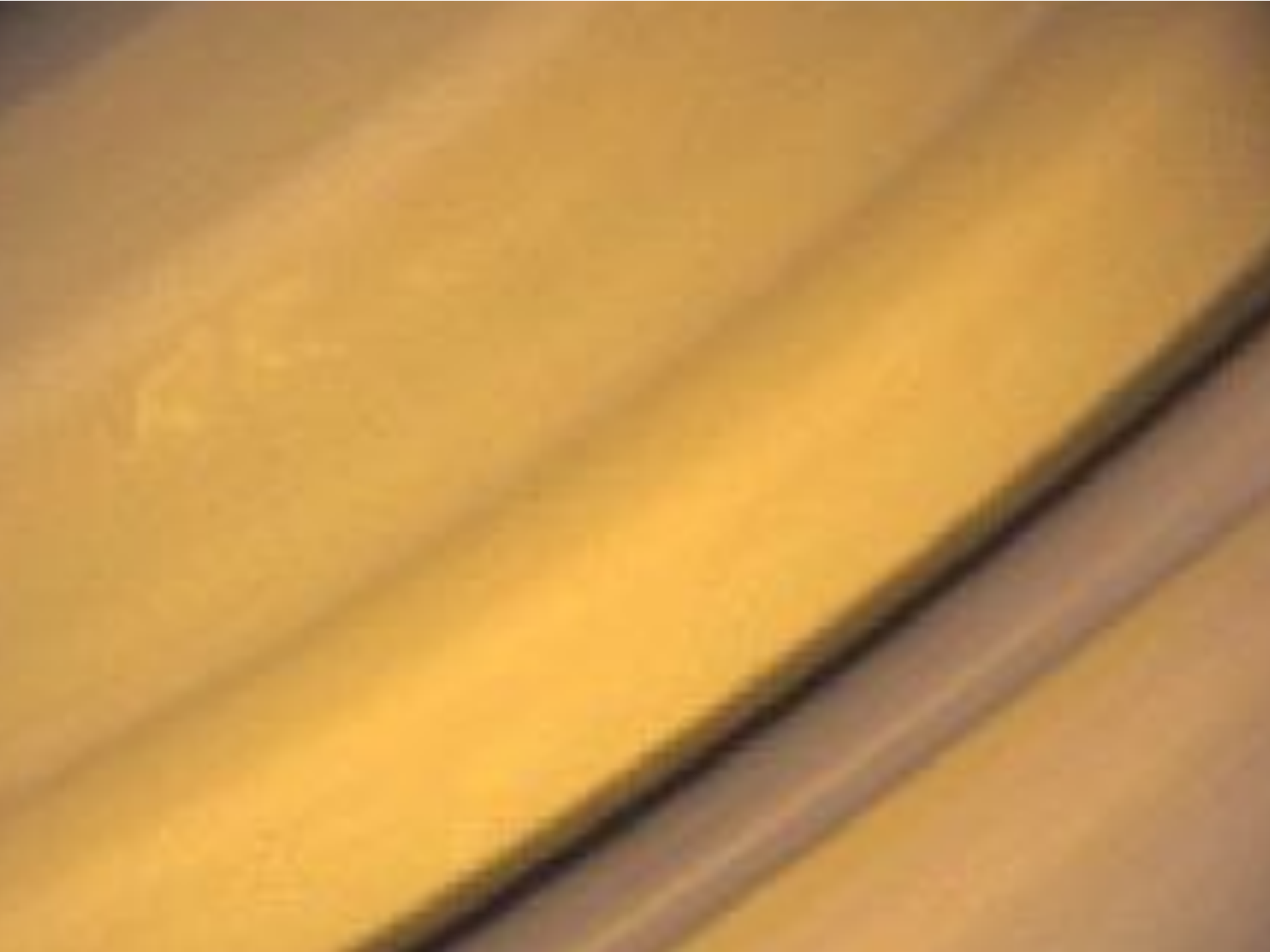
- ← T126 Cassini footprint : 6 km











END